Question 1 (i)
So we have 2D, N atoms, and 2N independent harmonic oscillators.

$$
\begin{aligned}
& Z_{1}=\sum^{n} e^{-\beta \hbar \omega\left(n+\frac{1}{2}\right)}=e^{-\frac{\beta \hbar \omega}{2}} \sum^{n} e^{-\beta \hbar \omega n}=\frac{1}{2} \operatorname{csch} \frac{\beta \hbar \omega}{2} \\
& Z_{2 N}=\left(Z_{1}\right)^{2 N}=\left(e^{\frac{\beta \hbar \omega}{2}}-e^{-\frac{\beta \hbar \omega}{2}}\right)^{-2 N} \\
& E=-\frac{\partial}{\partial \beta} \ln Z=2 N \frac{\partial}{\partial \beta}\left[\ln \left(2 \sinh \frac{\beta \hbar \omega}{2}\right)\right]=N \hbar \omega \operatorname{coth} \frac{\beta \hbar \omega}{2}
\end{aligned}
$$

Hint: be familiar with the conversions between hyperbolic and exponential functions.

## Question 1 (ii)

$$
C=\frac{\partial E}{\partial T}=-N \hbar \omega \operatorname{csch}^{2}\left(\frac{\beta \hbar \omega}{2}\right)\left(-\frac{\hbar \omega}{2 k T^{2}}\right)=\frac{N \hbar^{2} \omega^{2}}{2 k T^{2}} \operatorname{csch}^{2} \frac{\beta \hbar \omega}{2}
$$

At high temperature,

$$
C \approx \frac{N \hbar^{2} \omega^{2}}{2 k T^{2}} \frac{4}{\beta^{2} \hbar^{2} \omega^{2}}=2 N k
$$

C is constant in accordance to Curie's Law.

At low temperature,

$$
C=\frac{N \hbar^{2} \omega^{2}}{2 k T^{2}} \frac{1}{\left(e^{\frac{\beta \hbar \omega}{2}}-e^{-\frac{\beta \hbar \omega}{2}}\right)^{2}} \approx \frac{N \hbar^{2} \omega^{2}}{2 k T^{2}} e^{-\beta \hbar \omega} \approx 0
$$

Hint 1: Curie's Law at high temperature includes the following 3 things: $M \propto \frac{1}{T}, \chi \propto \frac{1}{T}$ and $C \propto k$.
Hint 2: $e^{-x} \approx 0$ when x is big, and $e^{x} \approx 1+x$ when x is close to zero. The graph of $y=\frac{e^{-\frac{1}{x}}}{x^{2}}$ looks something like the following, and it gets to zero as x tends to zero:


Question 1 (iii)

$$
S=k \ln Z-\frac{U}{T}=2 N k \ln \left(2 \sinh \frac{\beta \hbar \omega}{2}\right)-\frac{N \hbar \omega}{T} \operatorname{coth} \frac{\beta \hbar \omega}{2}
$$

For low temperature,

$$
S \approx 2 N k \ln e^{\frac{\beta \hbar \omega}{2}}-\frac{N \hbar \omega}{T}=\frac{2 N k \hbar \omega}{2 k T}-\frac{N \hbar \omega}{T}=0
$$

In accordance to the $3^{\text {rd }}$ Law of Thermodynamics.

At high temperature,

$$
S \approx 2 N k \ln \beta \hbar \omega-\frac{N \hbar \omega}{T} \frac{2}{\beta \hbar \omega}=2 N k\left[\ln \left(\frac{\hbar \omega}{k T}\right)-1\right]
$$

## Question 2 (i)

For $T \rightarrow 0, n_{k} \rightarrow \theta\left(\mu-\epsilon_{k}\right)$, where

$$
\begin{aligned}
& \theta\left(\mu-\epsilon_{k}\right)= \begin{cases}0, & \mu<\epsilon_{k} \\
1, & \mu>\epsilon_{k}\end{cases} \\
& N=\sum_{\epsilon<\epsilon_{k}} 1=\frac{2 A}{(2 \pi)^{2}} \pi k_{F}^{2}=\frac{A}{2 \pi} k_{F}^{2} \\
& k_{F}=\sqrt{\frac{2 \pi N}{A}} \\
& \epsilon_{F}=\frac{\hbar^{2} k_{F}^{2}}{2 m}=\frac{\hbar^{2}}{2 m}\left(2 \pi \frac{N}{A}\right)=\frac{\pi \hbar^{2}}{m} \frac{N}{A} \\
& U=\sum_{k<k_{F}} \epsilon_{k}=\frac{2 A}{(2 \pi)^{2}} \int_{0}^{k_{F}} \frac{\hbar^{2} k^{2}}{2 m} 2 \pi k d k=\frac{A \hbar^{2} k_{F}^{4}}{8 m \pi}=\frac{1}{2} N \epsilon_{F} \quad[\text { shown ] }
\end{aligned}
$$

Hint: Try thinking of $\theta\left(\mu-\epsilon_{k}\right)$ as a 2D Dirac delta function, a circle on a 2D plane with the origin as its centre. Another thing here is, $k_{F}$ is the wave factor, not to be confused with the Boltzmann constant!

## Question 2 (ii)

$$
N=\sum_{r} n_{r}=\frac{2 A}{(2 \pi)^{2}} \int \frac{2 \pi k d k}{e^{\beta(\epsilon-\mu)}+1}=\frac{m A}{\pi \hbar^{2}} \int \frac{m A}{\pi \hbar^{2}} \frac{d \epsilon}{e^{\beta(\epsilon-\mu)}+1}=\int n(\epsilon) f(\epsilon) d \epsilon
$$

So we can see that

$$
f(\epsilon)=\frac{A m}{\pi \hbar^{2}}=G^{\prime}(\epsilon), \quad G(\epsilon)=\frac{A m}{\pi \hbar^{2}} \epsilon, \quad G^{\prime \prime}(\epsilon)=0
$$

$$
\therefore N=\frac{A m}{\pi \hbar^{2}} \mu
$$

At $T=0$,

$$
N=\frac{m A}{\pi \hbar^{2}} \epsilon_{F}=\frac{m A}{\pi \hbar^{2}} k T_{F}
$$

$\therefore \mu=k T_{F}=\epsilon_{F}$, it is independent of temperature.

## Question 2 (iii)

$$
\begin{aligned}
& U=\int n(\epsilon) f(\epsilon) \epsilon d \epsilon=\int n(\epsilon) G^{\prime}(\epsilon) d \epsilon \\
& G^{\prime}(\epsilon)=\frac{A m}{\pi \hbar^{2}} \epsilon, \quad G(\epsilon)=\frac{A m}{2 \pi \hbar^{2}} \epsilon^{2}, \quad G^{\prime \prime}(\epsilon)=\frac{A m}{\pi \hbar^{2}} \\
& U=\frac{A m \mu^{2}}{2 \pi \hbar^{2}}+\frac{\pi^{2}}{6}(k T)^{2} \frac{A m}{\pi \hbar^{2}}=\frac{1}{2} N \mu+\frac{N \pi^{2}}{6} \frac{(k T)^{2}}{\mu}=\frac{1}{2} N k T_{F}\left[1+\frac{\pi^{2}}{3}\left(\frac{T}{T_{F}}\right)^{2}\right] \\
& C_{v}=\frac{\partial U}{\partial T}=\frac{1}{2} N k T_{F}\left(\frac{2 \pi^{2} T}{3 T_{F}^{2}}\right)=\frac{N k \pi^{2}}{3} \frac{T}{T_{F}}
\end{aligned}
$$

## Question 3 (i)

$n \lambda^{3}=e^{\mu \beta} \mp \frac{e^{2 \mu \beta}}{2^{\frac{3}{2}}}+\cdots$
$1^{\text {st }}$ order approximation, $n \lambda^{3}=e^{\mu \beta}$.
$2^{\text {nd }}$ order approximation, $n \lambda^{3}=e^{\mu \beta} \mp \frac{\left(n \lambda^{3}\right)^{2}}{2^{\frac{3}{2}}}$

$$
\begin{aligned}
& \Rightarrow e^{\mu \beta}=n \lambda^{3}\left(1 \pm \frac{n \lambda^{3}}{2^{\frac{3}{2}}}\right) \\
& \mu=k T\left[\ln n \lambda^{3}+\ln \left(1 \pm \frac{n \lambda^{3}}{2^{\frac{3}{2}}}\right)\right]
\end{aligned}
$$

For high temperature, $\lambda^{3}$ is small, so $\ln \left(1 \pm \frac{n \lambda^{3}}{2^{\frac{3}{2}}}\right) \approx \pm \frac{n \lambda^{3}}{2^{\frac{3}{2}}}$, and

$$
\therefore \mu=k T\left[\ln n \lambda^{3} \pm \frac{n \lambda^{3}}{2^{\frac{3}{2}}}+\cdots\right] \quad[\text { shown }]
$$

## Question 3 (ii)

$$
\begin{aligned}
& F=-k T \ln Z_{N}=-k T \ln \left[\frac{1}{N!}\left(\frac{V}{\lambda^{3}}\right)^{N}\right] \\
& P V=\frac{V}{\beta} \frac{\partial}{\partial V}\left(\ln Z_{N}\right)=k T V \frac{\partial}{\partial V}\left\{\ln \left[\frac{1}{N!}\left(\frac{V}{\lambda^{3}}\right)^{N}\right]\right\}=N k T \\
& G=N \mu=F+P V=k T\left[\ln N!-N \ln \left(\frac{V}{\lambda^{3}}\right)+N\right] \approx N k T\left[\ln N-\ln \left(\frac{V}{\lambda^{3}}\right)\right] \\
& \mu=k T \ln \left(\frac{N \lambda^{3}}{V}\right)=k T \ln \left(n \lambda^{3}\right) \quad[\text { shown }]
\end{aligned}
$$

Hint: Use Stirling's formula for the simplification of G.

## Question 3 (iii)



The fermions follow the blue line, the bosons the red line, while classical particles follow the green line. The intersection with the green line and the T axis is $T_{c}$. The quantum correction is the second term of the equation for $\mu$. The results show that the chemical potential curve for fermions is higher than classical particles, and it starts from a non-zero value of $\mu=\epsilon_{F}$. The boson curve is lower than classical and is always negative.

## Question 4 (i)

For the bose gas,

$$
\frac{P \lambda^{3}}{k T}=g_{\frac{5}{2}}(\zeta), \quad \frac{N \lambda^{3}}{V}=g_{\frac{3}{2}}(\zeta)
$$

At $T>T_{c}$,

$$
U=\frac{3}{2} P V=\frac{V}{\lambda^{3}} k T g_{\frac{5}{2}}(\zeta)=\frac{3}{2} \frac{k T}{\lambda^{3}} V g_{\frac{5}{2}}(\zeta)
$$

$$
\begin{aligned}
\Omega_{G} & =-P V=-\frac{k T}{\lambda^{3}} V g_{\frac{5}{2}}(\zeta) \\
G & =N \mu=N k T \ln \zeta \\
F & =\Omega_{G}+G=-\frac{k T}{\lambda^{3}} V g_{\frac{5}{2}}(\zeta)+N k T \ln \zeta \\
S & =\frac{U}{T}-\frac{F}{T}=\frac{3}{2} \frac{k}{\lambda^{3}} V g_{\frac{5}{2}}(\zeta)+\frac{k}{\lambda^{3}} V g_{\frac{5}{2}}(\zeta)-N k \ln \zeta=\frac{5}{2} \frac{k V}{\lambda^{3}} g_{\frac{5}{2}}(\zeta)-N k \ln \zeta \\
C_{v} & =\left(\frac{\partial U}{\partial T}\right)_{\beta \mu}=\frac{3}{2} k V \frac{\partial}{\partial T}\left[\frac{T}{\lambda^{3}} g_{\frac{5}{2}}(\zeta)\right]=\frac{5}{2} \frac{3}{2} \frac{k V}{\lambda^{3}} g_{\frac{5}{2}}(\zeta)+\frac{3}{2} \frac{k T V}{\lambda^{3}} \frac{\partial \zeta}{\partial T} \frac{\partial}{\partial \zeta} g_{\frac{5}{2}}(\zeta) \\
& =\frac{15}{4} \frac{V}{\lambda^{3}} g_{\frac{5}{2}}(\zeta)-\frac{9}{4} N k \frac{g_{\frac{3}{2}}(\zeta)}{g_{\frac{1}{2}}(\zeta)} \quad[\text { shown }]
\end{aligned}
$$

Hint: to find the expression of $C_{v}$, use the relation $\frac{\partial \zeta}{\partial T}=-\frac{3 N \lambda^{3}}{2 T V} \frac{\zeta}{g_{\overline{1}}(\zeta)}$, which can be obtained by partial differentiating both sides of the equation $\frac{N \lambda^{3}}{V}=g_{\frac{3}{2}}(\zeta)$ with respect to T .

## Question 4 (ii)

For $T<T_{c}$,

$$
\begin{aligned}
& \zeta=1, \quad g_{\frac{5}{2}}(1)=1.342, \quad g_{\frac{3}{2}}(1)=2.612, \quad g_{\frac{1}{2}}(1)=\infty \\
& U=\frac{3}{2} \frac{k T V}{\lambda^{3}}(1.342), \\
& \Omega_{G}=-\frac{1.342 k T V}{\lambda^{3}} \\
& G=0 \\
& F=\frac{1.342 k T V}{\lambda^{3}}=-\Omega_{G} \\
& S=\frac{5}{2} \frac{k V}{\lambda^{3}}(1.342) \\
& C_{v}=\frac{15}{4} \frac{V k}{\lambda^{3}}(1.342)
\end{aligned}
$$

## Question 4 (iii)

Since $T>T_{c}$,

$$
C_{v}=\frac{15}{4} \frac{V}{\lambda^{3}} g_{\frac{5}{2}}(\zeta)-\frac{9}{4} N k \frac{g_{\frac{3}{2}}(\zeta)}{g_{\frac{1}{2}}(\zeta)}=\frac{15}{4} N k \frac{g_{\frac{5}{2}}(\zeta)}{g_{\frac{3}{2}}(\zeta)}-\frac{9}{4} N k \frac{g_{\frac{3}{2}}(\zeta)}{g_{\frac{1}{2}}(\zeta)}
$$

At $T \gg T_{c}$,
$e^{\frac{\mu}{k T}} \rightarrow \frac{\mu}{k T}, \quad g_{k}(\zeta) \approx \zeta \approx n \lambda^{3}$
$\frac{C_{v}}{N k} \approx \frac{15}{4}-\frac{9}{4}=\frac{3}{2}$

At $T=T_{c}, \zeta=1$,
$\frac{C_{v}}{N k}=\frac{15}{4} \frac{1.342}{2.612} \approx 1.93$

At $T=0, C_{v}=0$. So our graph look like this:


Hint: Just in case you are confused, here's a table to summarize what equations are valid at what temperatures. In simple words, the two equations are valid above $T_{c}, \zeta=1$ at $T=T_{c}$, and the equation involving n is not valid below $T_{c}$ !

| Temperature <br> condition | The two equations should be |  |
| :---: | :--- | :--- |
| $T>T_{c}$ | $\frac{P \lambda^{3}}{k T}=g_{\frac{5}{2}}(\zeta)$, | $\frac{N \lambda^{3}}{V}=g_{\frac{3}{2}}(\zeta)$ |
| $T=T_{c}$ | $\frac{P \lambda^{3}}{k T}=g_{\frac{5}{2}}(1)$, | $\frac{N \lambda^{3}}{V}=g_{\frac{3}{2}}(1)$ |
| $T<T_{c}$ | $\frac{P \lambda^{3}}{k T}=g_{\frac{5}{2}}(1)$, | $\frac{N \lambda^{3}}{V}>g_{\frac{3}{2}}(1)$ |

Solutions provided by:
John Soo
© 2013, NUS Physics Society

